The iterations of intersection body operator.

Artem Zvavitch

Kent State University

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K is a star body if ρ_K(ξ) is positive and continuous function on Sⁿ⁻¹.
ξ[⊥] = {x ∈ ℝⁿ : x · ξ = 0}.

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Intersection Body

E. Lutwak: Intersection body, of a body K



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Intersection Body





R. Gardner, G. Zhang: More general definition: L is intersection body if it is limit in radial metric of IK.



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$$B_2^n = \{x \in \mathbb{R}^n : |x| \le 1\}$$
, then $IB_2^n = \operatorname{Vol}_{n-1}(B_2^{n-1})B_2^n = c_n B_2^n$.



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- $B_2^n = \{x \in \mathbb{R}^n : |x| \le 1\}$, then $IB_2^n = \operatorname{Vol}_{n-1}(B_2^{n-1})B_2^n = c_n B_2^n$.
- R. Gardner, A. Koldobsky, T. Schlumprecht: All convex symmetric bodies are intersection bodies in ℝⁿ, n ≤ 4. NOT true for n ≥ 5.

Spherical coordinates in ξ^{\perp}

$$\rho_{\mathrm{I}K}(\xi) = \mathrm{Vol}_{n-1}(K \cap \xi^{\perp}) = \frac{1}{n-1} \int_{S^{n-1} \cap \xi^{\perp}} \rho_K^{n-1}(\theta) d\theta = \frac{1}{n-1} R \rho_K^{n-1}(\xi).$$

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More general definition of Intersection Body (C^{∞} -case).

A symmetric star body *L* is an intersection body if $R^{-1}\rho_L \ge 0$.

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Consider body K such that for every $u \in S_{n-1}$ there exits an intersection body K_u , which coincide with K on a ε -neighborhood of u. Is it true that K must be an intersection body itself?

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Radon Transform: Fix $\varepsilon \in (0, 1/10)$

Consider a symmetric function f on S^{n-1} , such that for every $u \in S_{n-1}$ there exits a function f_u , which is equal to f on a ε -neighborhood of u and $R^{-1}f_u > 0$. Is it true that $R^{-1}f > 0$?

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F. Nazarov, D. Ryabogin, A. Z., 2008:

• NO!

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• NO!

• If we instead of regular neighborhoods around points would take neighborhood around equators then YES for even *n* and NO for odd *n*!!!

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- $d_{BM}(ITK, ITL) = d_{BM}(IK, IL).$
- $d_{BM}(B_2^n, \mathrm{I}B_2^n) = 1.$

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- $d_{BM}(K, IK) = 1, K \subset \mathbb{R}^2, K$ -symmetric.

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Examples:

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A. Fish, F. Nazarov, D. Ryabogin, A.Z.:

Consider a star body $K \subset \mathbb{R}^n$, $n \ge 3$, is it true that

$$d_{BM}(I^mK,B_2^n) \rightarrow 1, \text{ as } m \rightarrow \infty,$$

i.e. iterations of intersection body operator of a star body K will converge to B_2^n in d_{BM} ?

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 ΠL – projection body of L:

 $h_{\Pi L}(\theta) = \operatorname{Vol}_{n-1}(L|\theta^{\perp}).$

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Examples:

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 ΠL – projection body of *L*:

$$h_{\Pi L}(\theta) = \operatorname{Vol}_{n-1}(L|\theta^{\perp}).$$

Examples:

- $\Pi B_2^n = c_n B_2^n.$
- $\Pi B_{\infty}^n = c_n B_{\infty}^n$, where $B_{\infty}^n = [-1,1]^n$.

Fixed point is NOT unique! W. Weil (71) described polytopes that satisfy this property. General case is still open.

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Is it true that $d_{BM}(I^m K, B_2^n) \rightarrow 1$, as $m \rightarrow \infty$?

A. Fish, F, Nazarov, D. Ryabogin, A.Z., (2009)

 $\exists \varepsilon_n > 0$ such that $\forall K \subset \mathbb{R}^n$ such that K-start body, $d_{BM}(K, B_2^n) < 1 + \varepsilon_n$, we get

 $d_{BM}(I^m K, B_2^n) \rightarrow 1$, as $m \rightarrow \infty$.

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Remarks:

• We do not assume convexity of K. Such an assumption will much simplify the proofs.

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- Busemann theorem: If K-convex symmetric, then IK is convex symmetric.
- Even if K is convex symmetric, then d_{BM}(K, B₂ⁿ) ≤ √n, which is very far from ε_n.

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- Not known for convex symmetric case!
- (J. Kim, V. Yaskin, A.Z.) Wrong without assumption of convexity! there is a star body (*p*-convex) K such that d_{BM}(IK, Bⁿ₂) >> d_{BM}(K, Bⁿ₂).

Spherical Radon Transform:

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Denote by
$$\mathcal{R} = \frac{1}{\operatorname{Vol}_{n-2}(S^{n-2})}R$$
, i.e. $\mathcal{R}1 = 1$.

Question: $(n \ge 3)$

Consider symmetric function $f: S^{n-1} \to \mathbb{R}^+$, such that $f = \mathcal{R}f^{n-1}$, is it true that then f = 1?

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 \mathcal{H}_k - space of Spherical Harmonics of degree k.

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 \mathcal{H}_k - space of Spherical Harmonics of degree k. H_k^f the projection of f to \mathcal{H}_k , so

$$f \sim \sum_{k \ge 0} H_k^f$$

(Note: f-symmetric, we need only even k.)

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(Note: f-symmetric, we need only even k.)

Assume that $n \geq 3$. If $H_k \in \mathcal{H}_k$, k-even, then

$$\mathcal{R}H_k(\xi)= \mathsf{v}_{n,k}H_k(\xi), ext{ for all } \xi\in S^{n-1},$$

where $v_{n,0} = 1$ and

$$v_{n,k} = \frac{1 \cdot 3 \cdots (k-1)}{(n-1)(n+1) \cdots (n+k-3)}.$$

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- $\mathcal{R}f = \mathcal{R}g$, then f = g.
- $\mathcal{R}f = f$, then f = 1 (o.k. f = const).

THE MAIN PROBLEM:

$$f \sim \sum_{k \ge 0} H_k^f \Rightarrow$$

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Formulas Exists: Clebsch–Gordan coefficients — but they are hard, not clear (to me!) how to use for this problem.

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 $f = 1 + \phi$, where ϕ is even with small L_{∞} norm, $\int_{S^{n-1}} \phi = 0$.

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$$\begin{split} f &= 1 + \phi \text{, where } \phi \text{ is even with small } L_{\infty} \text{ norm, } \int_{S^{n-1}} \phi = 0. \\ \mathcal{R}f^{n-1} &= 1 + (n-1)\mathcal{R}\phi + \mathcal{R}O(\phi^2) \end{split}$$

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Problems:

1) Working with Spherical Harmonics we need to talk about L_2 norm! If we assume convexity, then those are "almost" equivalent. Much more work required to "prepare" the function to be ready for the L_2 , L_∞ game.

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Indeed, then $\|\mathcal{R}\phi^2\|_{L_2} \leq \|\phi\|_{L_\infty} \|\phi\|_{L_2}$ (do not forget $\|\mathcal{R}\|_{L_2 \to L_2} \leq 1$).

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1) Working with Spherical Harmonics we need to talk about L_2 norm! If we assume convexity, then those are "almost" equivalent. Much more work required to "prepare" the function to be ready for the L_2 , L_∞ game. 2) Our main goal to show that

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- If $f \in \mathcal{U}_{\alpha}$, then $\mathcal{R}f \in \mathcal{U}_{\alpha+n-2}$ and $\|\mathcal{R}f\|_{\mathcal{U}_{\alpha+n-2}} \leq C \|f\|_{\mathcal{U}_{\alpha}}$.

Make ρ_{I^kK} smooth!

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Let $\mu = \int f_k$. If $\varepsilon > 0$ is sufficiently small, then $|\mu - 1|$ is small and $\mu^{-1}f_k = 1 + \psi$ where $\int \psi = 0$ and $\|\psi\|_{L^{\infty}}$ is small. Note that

$$\|\psi\|_{\mathcal{U}_{\beta}} \leq 1 + \mu^{-1} \|f_k\|_{\mathcal{U}_{\beta}} \leq C'(k),$$

by (3), $\|\psi\|_{\mathcal{U}_{\alpha}}$ is also small $(\|\psi\|_{\mathcal{U}_{\beta}} < C(k) \text{ and } \|\psi\|_{L^{\infty}} \to 0 \text{ as } \varepsilon \to 0).$

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by (3), $\|\psi\|_{\mathcal{U}_{\alpha}}$ is also small $(\|\psi\|_{\mathcal{U}_{\beta}} < C(k) \text{ and } \|\psi\|_{L^{\infty}} \to 0 \text{ as } \varepsilon \to 0)$. Applying this to the function ρ_{K} , we conclude that if K is sufficiently close to B_{n} , then, after proper normalization, $\rho_{I^{k}K}$ can be written as $1 + \varphi$ with $\|\varphi\|_{\mathcal{U}_{\alpha}}$ as small as we want,